

## Examination paper for FY3201/FY8902 Atmospheric physics and climate change

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**Examination date: 9 June 2023**

**Examination time (from-to): 09:00-13:00**

**Permitted examination support material: Code G:**

Textbook (or printed pdf of textbook) with no annotations inside

One SIDE of A5 paper with handwritten or printed notes

ALL calculators are allowed.

**Other information:**

**Language: English**

**Number of pages (front page excluded): 7 pages**

**Number of pages enclosed: 7 pages (plus cover sheet)**

**Informasjon om trykking av eksamensoppgave**

**Originalen er:**

**1-sidig** ☒ **2-sidig** ☐

**sort/hvit** ☐ **farger** ☒

**skal ha flervalgskjema** ☐

**Checked by:**

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Date

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Signature

For multiple choice questions, each incorrect or blank answer will score zero points. Answers have been randomized and are not exact. You must choose the best answer. For all calculations use SI units!

You may take:

Molar mass of dry air:  $\sim 29 \text{ kg/kmole}$

Molar mass of helium:  $\sim 4 \text{ kg/kmole}$

Molar mass of  $\text{H}_2\text{O}$ :  $\sim 18 \text{ kg/kmole}$

Molar mass of  $\text{CO}_2$ :  $\sim 44 \text{ kg/kmole}$

$273 \text{ K} = 0^\circ\text{C}$     $1 \text{ hPa} = 10^2 \text{ Pa} = 10^2 \text{ N m}^{-2}$     $1 \text{ atm} = 1013 \text{ hPa}$     $g = 9.8 \text{ m s}^{-2}$  constant in  $z$     $c = 3 \times 10^8 \text{ m s}^{-1}$

Avagadro's number:  $N_A = 6.02 \times 10^{23} \text{ molecules/mole}$

Boltzmann's constant  $k = 1.38 \times 10^{-23} \text{ J/K}$

Stefan-Boltzmann constant:  $\sigma = 5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$

Planck Constant:  $h = 6.63 \times 10^{-34} \text{ J s}$

Solar photospheric temperature,  $T_s = 5786 \text{ K}$

Radius of the Sun =  $695800 \text{ km}$

Radius of the Earth =  $6370 \text{ km}$

$1 \text{ AU}$  (Earth-Sun distance) =  $150 \times 10^6 \text{ km}$

Radius of Venus =  $6051 \text{ km}$

Venus-Sun distance =  $0.72 \text{ AU}$

Radius of Mars =  $3396 \text{ km}$

Mars-Sun distance =  $1.52 \text{ AU}$

Latent heat of vaporization water:  $L_v = 2.5 \times 10^6 \text{ J kg}^{-1}$

Density of liquid water =  $1000 \text{ kg m}^{-3}$

Latent heat of sublimation ice:  $L_i = 2.8 \times 10^6 \text{ J kg}^{-1}$

Density of water vapour =  $5 \times 10^{-3} \text{ kg m}^{-3}$

Gas constant for water vapour:  $R_v = 461 \text{ J K}^{-1} \text{ kg}^{-1}$

Surface tension of water droplet  $75 \times 10^{-3} \text{ N m}^{-1}$

Values for dry air:  $C_p = 1004 \text{ J K}^{-1} \text{ kg}^{-1}$

$C_v = 718 \text{ J K}^{-1} \text{ kg}^{-1}$

$R_d = 287 \text{ J K}^{-1} \text{ kg}^{-1}$

$\gamma = C_p / C_v$

$\kappa = R_d / C_p$

$R_d = C_p - C_v$

$\Gamma_{\text{dair}} = 9.8 \text{ K/km}$

Clausius-Clapeyron relation:  $e_s = 6.112 \text{ hPa} \cdot \exp \left[ \frac{L_v}{R_v} \cdot \left( \frac{1}{273 \text{ K}} - \frac{1}{T} \right) \right]$

Some integrals that may be of use:

$$\int x^m e^{(ax)} dx = \frac{x^m e^{(ax)}}{a} - \frac{m \int x^{(m-1)} e^{(ax)} dx}{a}$$

$$\int x e^{(ax)} dx = \frac{e^{(ax)} (ax - 1)}{a^2}$$

$$\text{For } a > 0 \quad \int_0^\infty e^{(-ax)} dx = \frac{1}{a}$$

$$\int_X^\infty e^{(-ax)} dx = \frac{e^{(-aX)}}{a}$$

$$\int \frac{1}{a + bx} dx = \frac{\ln(a + bx)}{b}$$

# PERIODIC TABLE OF ELEMENTS

1 | 2 | 3 | 4 | 5 | 6 | 7 | 8 | 9 | 10 | 11 | 12 | 13 | 14 | 15 | 16 | 17 | 18

Atomic # Symbol		Name Weight		C		Solid		Liquid		Gas		Unknown		Metals		Metalloids		Nonmetals		Pnictogens		Chalcogens		Halogens		Noble gases																																																																																																																																																																																																																																																																											
1	H	Hydrogen	1.008	4	Be	Beryllium	9.0122	12	Mg	Magnesium	24.305	19	K	Potassium	39.098	20	Ca	Calcium	40.078	21	Sc	Scandium	44.956	22	Ti	Titanium	47.867	23	V	Vanadium	50.942	24	Cr	Chromium	51.996	25	Mn	Manganese	54.938	26	Fe	Iron	55.845	27	Co	Cobalt	58.933	28	Ni	Nickel	58.693	29	Cu	Copper	63.546	30	Zn	Zinc	65.38	31	Ga	Gallium	69.723	32	Ge	Germanium	72.630	33	As	Arsenic	74.922	34	Se	Selenium	78.971	35	Br	Bromine	79.904	36	Kr	Krypton	83.798	37	Rb	Rubidium	85.468	38	Sr	Strontium	87.62	39	Y	Yttrium	88.906	40	Zr	Zirconium	91.224	41	Nb	Niobium	92.906	42	Mo	Molybdenum	95.95	43	Tc	Technetium	(98)	44	Ru	Ruthenium	101.07	45	Rh	Rhodium	102.91	46	Pd	Palladium	106.42	47	Ag	Silver	107.87	48	Cd	Cadmium	112.41	49	In	Indium	114.82	50	Sn	Tin	118.71	51	Sb	Antimony	121.76	52	Te	Tellurium	127.60	53	I	Iodine	126.90	54	Xe	Xenon	131.29	55	Cs	Caesium	132.91	56	Ba	Barium	137.33	57–71	72	Hf	Hafnium	178.49	73	Ta	Tantalum	180.95	74	W	Tungsten	183.84	75	Re	Rhenium	186.21	76	Os	Osmium	190.23	77	Ir	Iridium	192.22	78	Pt	Platinum	195.08	79	Au	Gold	196.97	80	Hg	Mercury	200.59	81	Tl	Thallium	204.38	82	Pb	Lead	207.2	83	Bi	Bismuth	208.98	84	Po	Polonium	(209)	85	At	Astatine	(210)	86	Rn	Radon	(222)	87	Fr	Francium	(223)	88	Ra	Radium	(226)	89–103	104	Rf	Rutherfordium	(267)	105	Db	Dubnium	(268)	106	Sg	Seaborgium	(269)	107	Bh	Bohrium	(270)	108	Hs	Hassium	(277)	109	Mt	Meitnerium	(278)	110	Ds	Darmstadtium	(281)	111	Rg	Roentgenium	(282)	112	Cn	Copernicium	(285)	113	Nh	Nihonium	(286)	114	Fl	Flerovium	(289)	115	Mc	Moscovium	(290)	116	Lv	Livermorium	(293)	117	Ts	Tennessine	(294)	118	Og	Oganesson	(294)

C Solid

Hg Liquid

H Gas

Rf Unknown

Alkali metals

Alkaline earth metals

Lanthanoids (Lanthanides)

Actinoids (Actinides)

Transition metals

Post-transition metals

Metalloids

Other nonmetals

Noble gases

For elements with no stable isotopes, the mass number of the isotope with the longest half-life is in parentheses.

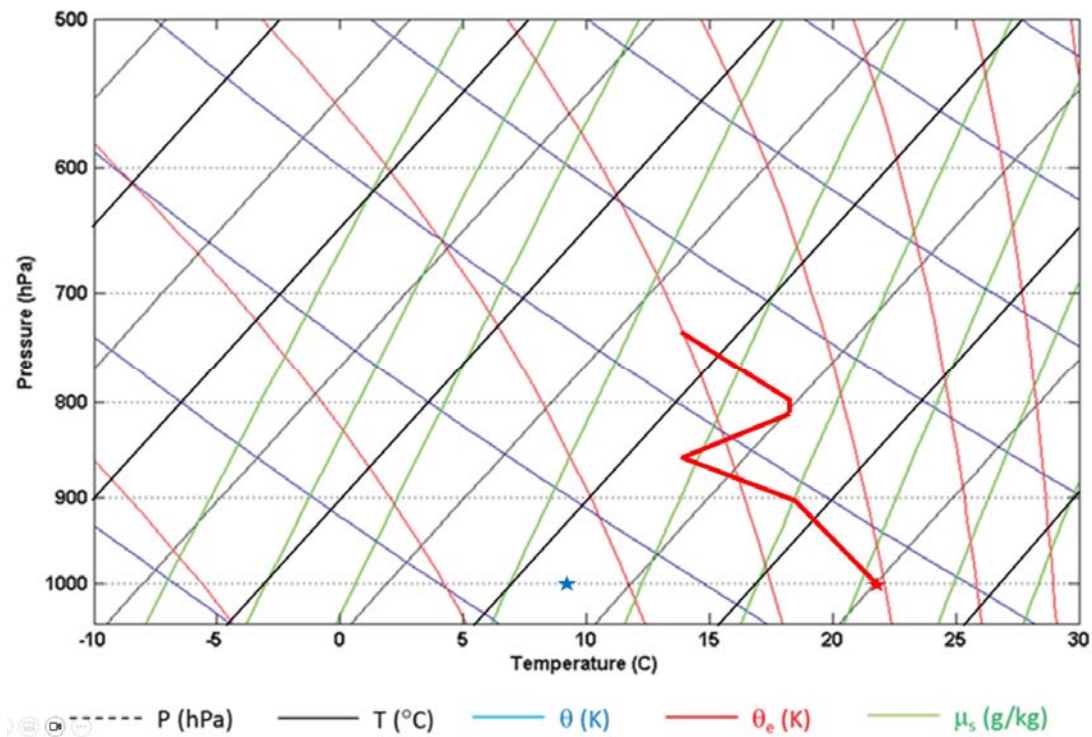
57 <b>La</b> Lanthanum 138.91	58 <b>Ce</b> Cerium 140.12	59 <b>Pr</b> Praseodymium 140.91	60 <b>Nd</b> Neodymium 144.24	61 <b>Pm</b> Promethium (145)	62 <b>Sm</b> Samarium 150.36	63 <b>Eu</b> Europium 151.96	64 <b>Gd</b> Gadolinium 157.25	65 <b>Tb</b> Terbium 158.93	66 <b>Dy</b> Dysprosium 162.50	67 <b>Ho</b> Holmium 164.93	68 <b>Er</b> Erbium 167.26	69 <b>Tm</b> Thulium 168.93	70 <b>Yb</b> Ytterbium 173.05	71 <b>Lu</b> Lutetium 174.97
89 <b>Ac</b> Actinium (227)	90 <b>Th</b> Thorium 232.04	91 <b>Pa</b> Protactinium 231.04	92 <b>U</b> Uranium 238.03	93 <b>Np</b> Neptunium (237)	94 <b>Pu</b> Plutonium (244)	95 <b>Am</b> Americium (243)	96 <b>Cm</b> Curium (247)	97 <b>Bk</b> Berkelium (247)	98 <b>Cf</b> Californium (251)	99 <b>Es</b> Einsteinium (252)	100 <b>Fm</b> Fermium (257)	101 <b>Md</b> Mendelevium (258)	102 <b>No</b> Nobelium (259)	103 <b>Lr</b> Lawrencium (266)



- 1) (5 pts) A parcel of air at 900 hPa has a temperature of 20 °C. It is lifted adiabatically until the parcel temperature drops to 13 °C. What is the pressure at this altitude?
- 2) (5 Pts) A parcel of air at 1000 hPa and 25 °C is lifted adiabatically to its lifting condensation level at 800 hPa where the parcel temperature is  $T=7$  °C, what was the relative humidity of the parcel before lifting?
- 3) (20 pts) In the balloon sounding on the SkewT-LnP diagram below, the atmospheric temperature is plotted in red as a function of pressure. At 1000 hPa is 20 °C and the dew point is 7 °C.
  - a) At what pressure is the lifting condensation level (LCL)?

State the stability conditions (stable, unstable, neutrally or conditionally stable) of the atmosphere to small vertical displacements in the following pressure ranges:

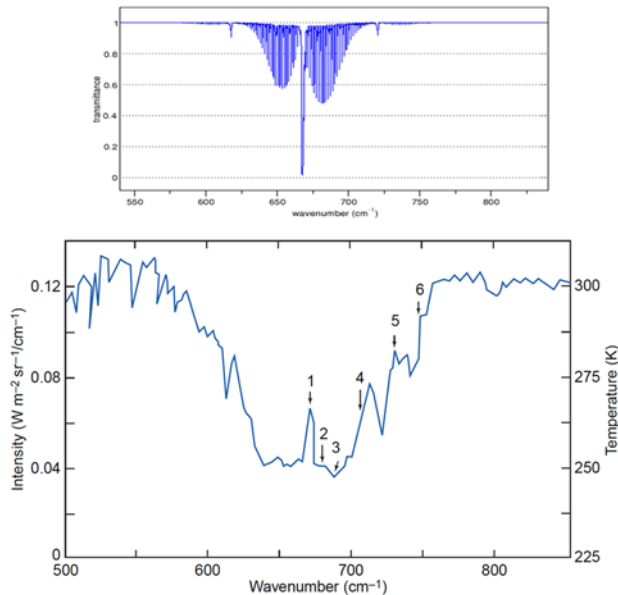
- b) Between 1000 and 900 hPa
- c) Between 900 and 850 hPa
- d) Between 850 and 800 hPa
- e) Between 800 and 750 hPa



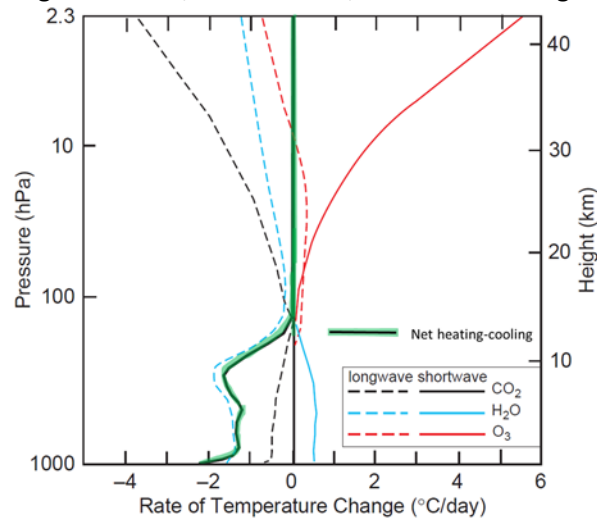
**Questions 4-8 are multiple choice, and no calculations need to be shown.**

- 4) (4 pts) The \_\_\_\_\_ and the \_\_\_\_\_ are layers of the atmosphere that can become unstable.
- a) troposphere, ionosphere
  - b) mesosphere, troposphere
  - c) stratosphere, thermosphere
  - d) mesosphere, cryosphere
- 5) (4 pts) What is the major species in the stratosphere?
- a) Molecular Oxygen
  - b) Atomic Oxygen
  - c) Molecular Nitrogen
  - d) Ozone
  - e) None of the above.
- 6) (4 pts) Sketch the relative spectral radiance as a function of wavelength for three blackbodies at temperatures  $T_1 > T_2 > T_3$ . Label the curves with their temperatures and put units on the axes, but you do not need to put numerical values on the axes.
- 7) (4 pts) Relative to the Earth's surface, what effect does the Coriolis force have on masses of air or water that are changing latitude?
- a) The results are unpredictable; currents can veer right or left in either hemisphere.
  - b) They turn to the left in the northern hemisphere and to the right in the southern hemisphere.
  - c) They turn to the right in the northern hemisphere and to the left in the southern hemisphere.
  - d) They turn to the right in both hemispheres.
  - e) They turn to the left in both hemispheres.
- 8) (4 pts) What wavelengths of sunlight are absorbed by molecular nitrogen in the troposphere?
- a) Infrared.
  - b) Ultraviolet.
  - c) Radio waves.
  - d) Microwaves.
  - e) Visible.
  - f) None of the above.

- 9) (5 Pts) Shown below is the carbon dioxide 15-micron band. The top figure is the unsaturated absorption spectrum taken in the laboratory, and below is the intensity of the outgoing radiation from Earth measured by satellite showing the carbon dioxide. Can you explain why the intensity (and temperature) of the outgoing radiation at point 1 is higher than that at points 2 and 3?



- 10) (5 Pts) Shown below are the vertical profiles of the time rate of change of temperature due to the absorption of solar radiation (solid curves) and the absorption and emission of infrared radiation (dashed curves) by water vapor (blue), carbon dioxide (black), and ozone (red). The heavy black/green solid curve represents the net heating or cooling effect of the three gases. The figure shows that the net effect of the absorption and emission of radiation by greenhouse gases is to cool the air in the troposphere about 1°C per day. Can you explain why an increase of greenhouse gases would, nevertheless, result in warming the air in the troposphere?



- 11) (5 Pts) What ozone column (in molecules·m<sup>-2</sup>) corresponds to 1 Dobson Unit (DU)?

12) Nitrogen dioxide, (NO<sub>2</sub>) is a pollutant released by diesel and petrol internal combustion engines as well as industrial processes that causes smog. Mainly produced by traffic in cities, the narrow valleys and fjords where Norwegian cities are located tend to concentrate levels. However, the yearly average mass density across Norway 32.33 µg·m<sup>-3</sup>, which is low by European standards.

- a) (5 Pts) What is the average number density of NO<sub>2</sub> in molecules·m<sup>-3</sup>?
- b) (5 Pts) At 1000 hPa and 22 C, the density of NO<sub>2</sub> measured one day in Bergen was 29.95 µg·m<sup>-3</sup> what is the volume mixing ratio of NO<sub>2</sub>?
- c) (5 pts) Originally, the Martian atmosphere was thought to consist only of 95% CO<sub>2</sub>, 3% N<sub>2</sub> and 2% Ar. However, recently, it was discovered that it also contains 0.2% O<sub>2</sub>. By how much did the initial estimate of the mean molecular weight of the Martin atmosphere change when this additional species, O<sub>2</sub> was taken into account.

13) On Venus  $g_v=8.87 \text{ m}\cdot\text{s}^{-2}$ , the specific gas constant for dry air is  $R_{dv} = 195.5 \text{ J}\cdot\text{K}^{-1}\cdot\text{kg}^{-1}$ , and the specific heat at constant pressure,  $C_{pv} = 846 \text{ J}\cdot\text{K}^{-1}\cdot\text{kg}^{-1}$ . The pressure at the surface is 92000 hPa, and the temperature 753 K. The temperature falls off linearly with a lapse rate of 8 K/km (note non-SI units of hPa and km).

- a) (2 Pts) What is the buoyancy period at 59 km, and is the atmosphere stable with respect to vertical motion there?
- b) (6 Pts) What is the temperature and pressure at 59 km?

The Venusian atmosphere is composed of 96.5% CO<sub>2</sub> and approximately 3.5% N<sub>2</sub>, neither of which absorb solar radiation. However, there is a 150-ppm mixing ratio of sulfur dioxide (SO<sub>2</sub>) at the surface that absorbs at in the UV/visible with an absorption coefficient  $k = 0.11 \text{ m}^2\cdot\text{kg}^{-1}$ . The Sun is at a solar **ZENITH** angle (angle of the Sun from overhead) of  $\chi = 45$  degrees. You may assume no scattering, and take the optical depth of incoming solar radiation (averaged over all wavelengths), the density of air, and the mass mixing ratio of the absorber to be defined as:

$$\tau(Z) = \int_Z^{\infty} \rho(z) \mu(z) k dz$$

$$\rho(z) = \rho_0 e^{\left(-\frac{z}{H_m}\right)}$$

$$\mu(z) = \mu_0 e^{\left(-\frac{z}{H_v}\right)}$$

Where  $\rho(z)$  is the mass density of air,  $\mu(z)$  is the mass mixing ratio of SO<sub>2</sub>.  $H_m$  is the scale height of the atmosphere is  $H_m = 13 \text{ km}$ , and the scale height of the SO<sub>2</sub>, is  $H_v = 8 \text{ km}$ . At the surface, the pressure is 92000 hPa, the total density,  $\rho_0=85 \text{ kg}\cdot\text{m}^{-3}$  and the temperature is 753 K

- c) (6 Pts) Calculate the optical depth  $\tau$  and its derivative with height  $d\tau/dz$  at a height of 5 km in the atmosphere.
- d) (6 Pts) Calculate the heating rate in  $\text{Kelvin}\cdot\text{day}^{-1}$  due to short wavelength radiation at 11 km if the downward solar flux at the top of the atmosphere is  $2600 \text{ W/m}^2$  and the solar ZENITH angle is  $\chi=45$  degrees. At 11 km, the total density is  $\rho=36.5 \text{ kg}\cdot\text{m}^{-3}$ , the absorber mass mixing ratio is  $0.0559 \text{ g}\cdot\text{kg}^{-1}$ , the optical depth is  $\tau=1$  and its derivative,  $d\tau/dz$ , is  $-0.22 \text{ km}^{-1}$ .