TFY4345 Classical Mechanics Suggested Solution to Exam November 27, 2023

Part I

- 1.1 B Such a constraint, expressed in terms of an equation, is called holonomic.
- **1.2 B** In the element ε_{321} of the Levi-Civita tensor, the indices come in a 'non-cyclic' order, but they are all different. Then the value is -1.
- **1.3** A An angle is dimensionless. Then, $p = \partial L/\partial \dot{\theta}$ has the unit Js, i.e., the unit of an angular momentum.
- 1.4 E A coordinate not present in the Lagrangian is called cyclic.
- **1.5 E** Here, the kinetic energy $T = m(\dot{x}^2 + \dot{y}^2)/2$ will not be a function of θ . Then p_{θ} is conserved.
- **1.6** D $n = 3N k = 3 \cdot 3 3 = 6$ independent coordinates.
- **1.7 B** We have 4 independent coordinates: x for the sliding support at A and an angle θ_j (j = 1, 2, 3) between (e.g.) the vertical axis and each massless rod.

1.8 C
$$T = (m_1 + m_2)\dot{x}^2/2$$
 and $V = -m_1gx - m_2g(\ell - x)$ gives

$$L = \frac{1}{2}(m_1 + m_2)\dot{x}^2 + (m_1 - m_2)gx + m_2g\ell$$

- **1.9** F $p_2 = \partial L/\partial \dot{x}_2 = m\dot{x}_2 + qA_2$.
- **1.10 B** $B_i = \varepsilon_{ijk} \partial_j A_k$. Here, $B_1 = B_2 = 0$ and $B_3 = \varepsilon_{312} \partial_1 A_2 + \varepsilon_{321} \partial_2 A_1 = B_0 (-B_0) = 2B_0$.
- **1.11** E $E = -\partial A/\partial t \nabla \phi = -E_0(\hat{x} + \hat{y} + \hat{z}).$
- **1.12** C $p_1 = \partial L/\partial \dot{q}_1 = c_3 q_1$.
- **1.13** C $r_{\min}/r_{\max} = [p/(1+\varepsilon)]/[p/(1-\varepsilon)] = (1-\varepsilon)/(1+\varepsilon) = 0.79/1.21 = 0.65.$
- **1.14 A** Compared with the 'usual' example of rotation around the x_3 axis, the 3 axis now plays the role of the 1 axis, and the 1 axis now plays the role of the 2 axis. Therefore, the rotation matrix is

$$\begin{pmatrix}
\cos\phi & 0 & -\sin\phi \\
0 & 1 & 0 \\
\sin\phi & 0 & \cos\phi
\end{pmatrix}$$

- **1.15** A $v = c \cdot (5/6 + 7/8)/(1 + 5 \cdot 7/6 \cdot 8) = 82c/83$.
- **1.16 D** In this question, and the next, the angular velocity of Earth is $\omega = 2\pi/T = [2\pi/(24 \cdot 3600)] \text{ s}^{-1}$. Then $a_{\text{Cor}} = 2v_b\omega = 2 \cdot 15 \cdot 2\pi/(24 \cdot 3600) \text{ m/s}^2 = 2.2 \text{ mm/s}^2$.
- **1.17 D** $a_{\text{Cen}} = \omega^2 r = (2\pi/(24 \cdot 3600))^2 \cdot 6.378 \cdot 10^6 \text{ m/s}^2 = 34 \text{ mm/s}^2$.

1.18 F We have $E^2 = (pc)^2 + (mc^2)^2$, i.e., $m = \sqrt{E^2 - (pc)^2}/c^2 = \sqrt{500^2 - 400^2}$ MeV/ $c^2 = 300$ MeV/ c^2 .

1.19 F $p = \partial F_1/\partial q = Q - 2q$, i.e., Q(q, p) = p + 2q. Then, $P = -\partial F_1/\partial Q = -q + 2Q = -q + 2 \cdot (p + 2q) = 3q + 2p$.

1.20 C A key ingredient in Unni's Cinnamon Cake is Cinnamon.

Part II

 $\mathbf{2}$

a) Stretch of left spring:

$$(x_2 - x_1) - (x_{20} - x_{10}) = \eta_2 - \eta_1$$

Stretch of right spring:

$$(x_3 - x_2) - (x_{30} - x_{20}) = \eta_3 - \eta_2$$

(Compression when $\eta_2 - \eta_1$ or $\eta_3 - \eta_2$ negative.) Potential energy:

$$V = \frac{1}{2}k(\eta_2 - \eta_1)^2 + \frac{1}{2}k(\eta_3 - \eta_2)^2$$

= $\frac{1}{2}k(\eta_1^2 + 2\eta_2^2 + \eta_3^2 - \eta_1\eta_2 - \eta_2\eta_1 - \eta_2\eta_3 - \eta_3\eta_2)$

The matrix V is:

$$\mathbf{V} = \left(\begin{array}{ccc} k & -k & 0 \\ -k & 2k & -k \\ 0 & -k & k \end{array} \right)$$

Kinetic energy:

$$T = \frac{1}{2} m_i \dot{x}_i^2 = \frac{1}{2} m_i \dot{\eta}_i^2$$

The matrix T is:

$$m{T} = \left(egin{array}{ccc} m & 0 & 0 \\ 0 & 2m & 0 \\ 0 & 0 & 4m \end{array}
ight)$$

b) The eigenfrequencies are given by the secular equation

$$\begin{vmatrix} k - \omega^2 m & -k & 0 \\ -k & 2k - 2\omega^2 m & -k \\ 0 & -k & k - 4\omega^2 m \end{vmatrix} = 0$$

If we extract a common factor k^3 and introduce the dimensionless factor $\alpha = m\omega^2/k$, the resulting equation is

$$\begin{vmatrix} 1 - \alpha & -1 & 0 \\ -1 & 2 - 2\alpha & -1 \\ 0 & -1 & 1 - 4\alpha \end{vmatrix} = 0$$

We expand the determinant by (e.g.) the upper row and find

$$(1 - \alpha)(2 - 2\alpha)(1 - 4\alpha) - (1 - \alpha) - (1 - 4\alpha) = 0$$

or

$$8\alpha^3 - 18\alpha^2 + 7\alpha = 0.$$

The normal mode with $\alpha = 0$ corresponds to pure translation of the system along the x axis. This is of no interest here. The two non-zero solutions are

$$\alpha = \frac{18 \pm \sqrt{18^2 - 4 \cdot 8 \cdot 7}}{2 \cdot 8} = \frac{18 \pm 10}{16} = \frac{9 \pm 5}{8},$$

i.e.,

$$\alpha_1 = \frac{1}{2} \quad , \quad \alpha_2 = \frac{7}{4}.$$

This corresponds to the frequencies

$$f_1 = \frac{\omega_1}{2\pi} = \frac{\sqrt{k\alpha_1/m}}{2\pi} = \frac{\sqrt{k/2m}}{2\pi} \simeq 6.2 \text{ Hz}$$

and

$$f_2 = \frac{\omega_2}{2\pi} = \frac{\sqrt{k\alpha_2/m}}{2\pi} = \frac{\sqrt{7k/4m}}{2\pi} \simeq 11.5 \text{ Hz}.$$

c) We insert the smallest value $\alpha = 1/2$ into the equation

$$\begin{pmatrix} 1 - \alpha & -1 & 0 \\ -1 & 2 - 2\alpha & -1 \\ 0 & -1 & 1 - 4\alpha \end{pmatrix} \begin{pmatrix} A_1 \\ A_2 \\ A_3 \end{pmatrix} = 0$$

and find

$$\frac{1}{2}A_1 - A_2 = 0$$

$$-A_1 + A_2 - A_3 = 0$$

$$-A_2 - A_3 = 0$$

Here, we have only two independent equations, so we can only find e.g. A_2 and A_3 expressed in terms of A_1 . If we set $A_1 = 1$, we find $A_2 = 1/2$ and $A_3 = -1/2$. In this normal mode, m and 2m oscillate in phase, and out of phase with 4m, with amplitudes that ensures a mass center at rest.

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a) Potential energy: $V = -mg\ell \cos \theta$. Kinetic energy (upper mass 1 and lower mass

Kinetic energy (upper mass 1 and lower mass 2):

$$T_1 = \frac{1}{2}m\dot{x}^2$$
 , $T_2 = \frac{1}{2}m(\dot{x}^2 + \dot{y}^2)$.

For the two masses, $x = \pm (\ell/2) \sin \theta$, and for the lower mass, $y = -\ell \cos \theta$. Then, $\dot{x} = \pm (\ell \dot{\theta}/2) \cos \theta$ and $\dot{y} = \ell \dot{\theta} \sin \theta$. We insert this into T_1 and T_2 and find, after using the hint,

$$L = T - V = T_1 + T_2 - V = \frac{m\ell^2\dot{\theta}^2}{4} \left(1 + \sin^2 \theta \right) + mg\ell \cos \theta.$$

b) For the equation of motion, we need these derivatives:

$$\frac{\partial L}{\partial \theta} = \frac{m\ell^2\dot{\theta}^2}{2}\sin\theta\cos\theta - mg\ell\sin\theta
\frac{\partial L}{\partial \dot{\theta}} = \frac{m\ell^2\dot{\theta}}{2}\left(1 + \sin^2\theta\right)
\frac{d}{dt}\frac{\partial L}{\partial \dot{\theta}} = \frac{m\ell^2\ddot{\theta}}{2}\left(1 + \sin^2\theta\right) + m\ell^2\dot{\theta}^2\sin\theta\cos\theta$$

We multiply the Lagrange equation with $2/m\ell^2$ and find

$$\ddot{\theta}(1+\sin^2\theta) + \sin\theta \left(\frac{2g}{\ell} + \dot{\theta}^2\cos\theta\right) = 0.$$

c) To linear order, $\sin \theta = \theta$ and $\cos \theta = 1$. Then,

$$\ddot{\theta} + \frac{2g}{\ell}\theta = 0.$$

This is a harmonic oscillator with frequency $\omega = \sqrt{2g/\ell}$.

d) We have already

$$p_{\theta} = \frac{\partial L}{\partial \dot{\theta}} = \frac{m\ell^2 \dot{\theta}}{2} \left(1 + \sin^2 \theta \right)$$

which means that

$$\dot{\theta} = \frac{2p_{\theta}}{m\ell^2(1+\sin^2\theta)}.$$

We insert this expression for $\dot{\theta}$ in L and find

$$H = p_{\theta}\dot{\theta} - L = \frac{p_{\theta}^2}{m\ell^2(1 + \sin^2\theta)} - mg\ell\cos\theta.$$

4 When the impact parameter is $s = a/\sqrt{2}$, the point particle collides with the surface of the disc with an angle $\pi/4$ between the incoming direction and the surface normal. The collision is specular, so the scattering angle is $\pi/2$.