

7. ^{theory of} Special relativity

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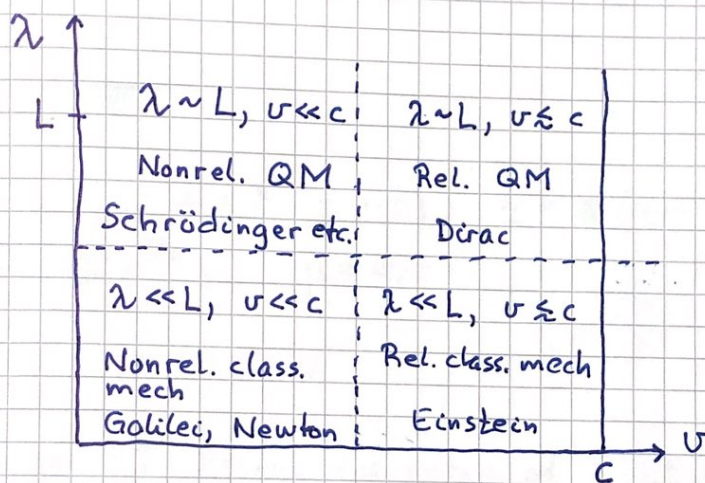
7.1 Introduction

L = linear size of system

$\lambda = h/p$ = de Broglie wavelength for particle (body)
(p = momentum, h = Planck constant)

v = typical particle velocity

c = speed of light



We will consider systems with $\lambda \ll L$ and $v \approx c$,
described by Einstein's special theory of relativity (SR)
[See GPS 7.11 for a very brief outline of the general theory (GR)]

Galilei: Assume two inertial systems S and S' ; S' moves with velocity $\vec{v}$ relative to S . A particle's coordinates $\vec{r}$ and $\vec{r}'$ and times t and t' observed in S and S' are related as $\vec{r}' = \vec{r} - \vec{v} \cdot t$ and $t' = t$. (GT)
The speed of light is different in S and S' , $c' = c - v$.

Experiments: $c' = c$ [Michelson - Morley etc]

Einstein's two postulates:

1. The laws of physics are the same in all inertial systems
2. The speed of light is _____ "

7.2 The Lorentz transformation (LT)

LT with vectors ($\vec{v} = c\vec{\beta}$):

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$$\vec{r}' = \vec{r} + (\gamma - 1) \frac{(\vec{\beta} \cdot \vec{r}) \vec{\beta}}{\beta^2} - \gamma c t \vec{\beta}$$

$$t' = \gamma t - \frac{\gamma}{c} \vec{\beta} \cdot \vec{r}$$

LT is linear $\Rightarrow$ may use ideas/results from chapter 4

Minkowski space:

4D space with axes $x_1=x, x_2=y, x_3=z$ and $x_4=ict$

$$\Rightarrow x^2 + y^2 + z^2 - c^2 t^2 = \sum_{\mu=1}^4 x_{\mu} x_{\mu} \equiv x_{\mu} x_{\mu} \quad [\text{sum conv.}]$$

Index notation: Greek for 4-vectors ($\alpha, \beta, \mu, \dots$)

Roman for 3-vectors ($i, j, k, \dots$)

$\Rightarrow$ Invariance of the speed of light can be expressed as

" $x_{\mu} x_{\mu}$ is an invariant"

Or: The norm (abs. value) of the "position vector" in Minkowski space is constant under a LT

From 4.2 (p.55): A linear transf. that leaves the norm of the vector unchanged is an orthogonal transf.

$\Rightarrow$ LT is an orthogonal transf. in Minkowski space

LT on matrix form:

$$x' = \mathbb{L} x \quad \text{or} \quad x'_{\mu} = L_{\mu\nu} x_{\nu}$$

With relative velocity along x_3 axis ($\vec{v} = v \hat{x}_3$):

$$\mathbb{L} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & \gamma & i\beta\gamma \\ 0 & 0 & -i\beta\gamma & \gamma \end{pmatrix}$$

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From 4.3 (p. 58) : $\mathbb{L}^{-1} = \tilde{\mathbb{L}}$ (= transposed of $\mathbb{L}$)

Inverse LT : $X = \mathbb{L}^{-1} X' = \tilde{\mathbb{L}} X' \Rightarrow x_\mu = L_{\nu\mu} x'_\nu$

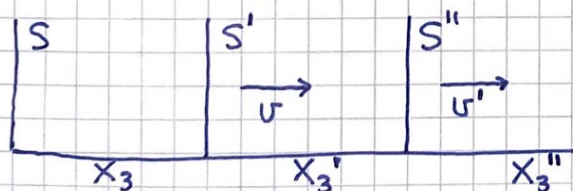
$$\mathbb{L}^{-1} = \tilde{\mathbb{L}} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & \gamma & -i\beta\gamma \\ 0 & 0 & i\beta\gamma & \gamma \end{pmatrix} \quad [v \rightarrow -v \text{ as seen before}]$$

Trigonometric convenience :

From example p. 56 : $\begin{pmatrix} \cos\varphi & \sin\varphi \\ -\sin\varphi & \cos\varphi \end{pmatrix}$ repr. rotation in a plane

Then, with $\gamma = \cos\varphi$ and $i\beta\gamma = \sin\varphi$, a LT can be viewed as a rotation an imaginary angle φ in the $x_3 x_4$ plane.

Ex : Velocity addition



v = speed of S' rel. to S

$v' = \text{---} \text{---} \text{---} S'' \text{---} \text{---} S'$

$v'' = \text{---} \text{---} \text{---} S'' \text{---} \text{---} S$

$v'' = ?$ GT : $v'' = v' + v$

$$X'' = \mathbb{L}' X' = \mathbb{L}' \mathbb{L} X = \mathbb{L}'' X$$

$$\Rightarrow L''_{33} = \cos\varphi' \cos\varphi - \sin\varphi' \sin\varphi = \cos(\varphi' + \varphi) = \cos\varphi''$$

$$L''_{34} = \cos\varphi' \sin\varphi + \sin\varphi' \cos\varphi = \sin(\varphi' + \varphi) = \sin\varphi''$$

$$\Rightarrow \tan\varphi'' = \frac{L''_{34}}{L''_{33}} = \dots = \frac{\tan\varphi + \tan\varphi'}{1 - \tan\varphi \tan\varphi'}$$

$$\Rightarrow i\beta'' = \frac{i\beta + i\beta'}{1 + \beta\beta'} \Rightarrow \underline{v'' = \frac{v + v'}{1 + vv'/c^2}} ; \text{Einstein's addition formula}$$

7.3 Covariant four-dimensional formulation

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We will address Einstein's 1. postulate (the principle of relativity):

Are the laws of physics the same in all inertial systems?

Or:

Are the equations (describing the laws of physics) covariant, i.e., invariant in form, under a LT?

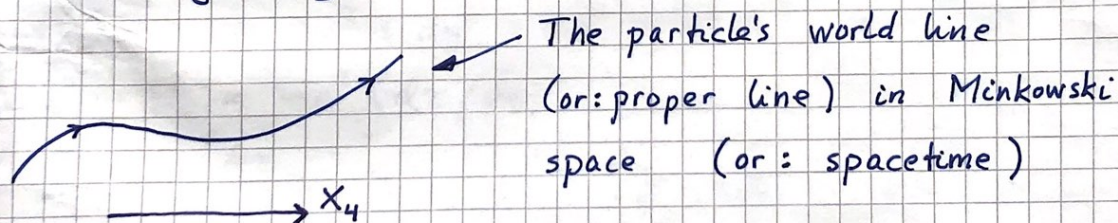
Laws-of-physics-equations provide the relation between scalars, vectors and tensors (expressed by matrices).

Scalars and vectors are special cases of tensors.

Hence, we must in general have covariance of tensor equations:

If $G_{\mu\nu} = D_{\mu\nu}$, then $G'_{\mu\nu} = D'_{\mu\nu}$ after a LT

Consider a particle (or observer), moving in 3D space as time goes by; $u^2 = (dx_i/dt) \cdot (dx_i/dt)$:



dx_μ = small element along the world line

Proper time τ : $dx_\mu dx_\mu \equiv -c^2 d\tau^2$ (def. of τ)

Let us show that

$$dx'_\mu dx'_\mu = dx_\mu dx_\mu$$

i.e., the same in S' and S , i.e., a Lorentz invariant

From 7.2 (and 4.3):

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$$X'^\mu = \mathbb{L}^\mu{}_\nu X^\nu; \quad X'_\mu = L_{\mu\nu} X^\nu$$

$$\mathbb{L}^{-1} = \tilde{\mathbb{L}} \Rightarrow \mathbb{L} \tilde{\mathbb{L}} = \tilde{\mathbb{L}} \mathbb{L} = \mathbb{1} \quad (\text{since } \mathbb{L} \mathbb{L}^{-1} = \mathbb{L}^{-1} \mathbb{L} = \mathbb{1})$$

Then:

$$\begin{aligned} dx'_\mu dx'_\mu &= L_{\mu\nu} dx_\nu \cdot L_{\mu\alpha} dx_\alpha \\ &= L_{\mu\nu} \tilde{L}_{\alpha\mu} dx_\nu dx_\alpha \\ &= \delta_{\alpha\nu} dx_\nu dx_\alpha = dx_\alpha dx_\alpha = dx_\mu dx_\mu \end{aligned}$$

$$\text{I.e., } dx'_\mu dx'_\mu = dx_\mu dx_\mu = -c^2 d\tau^2$$

Time dilation:

Assume particle at rest in S' ; $u=v$ = speed of S' rehto S

$$\Rightarrow dx'_\mu = (0, 0, 0, ic dt')$$

$$\Rightarrow dx'_\mu dx'_\mu = -c^2 dt'^2 = -c^2 d\tau^2$$

$$\Rightarrow d\tau = dt' = \text{time measured on clock at rest}$$

$$\text{Measured in } S: dx_\mu = (dx, dy, dz, ic dt)$$

$$\Rightarrow dx_\mu dx_\mu = dx^2 + dy^2 + dz^2 - c^2 dt^2 = -c^2 d\tau^2$$

$$\Rightarrow u^2 - c^2 = -c^2 (d\tau/dt)^2$$

$$\Rightarrow dt = \frac{d\tau}{\sqrt{1-\beta^2}} > d\tau \quad (\beta = u/c = v/c)$$

$$\Rightarrow dt = \gamma d\tau = \text{time measured on a clock in } S,$$

where the particle ~~moves~~ moves with speed v ,

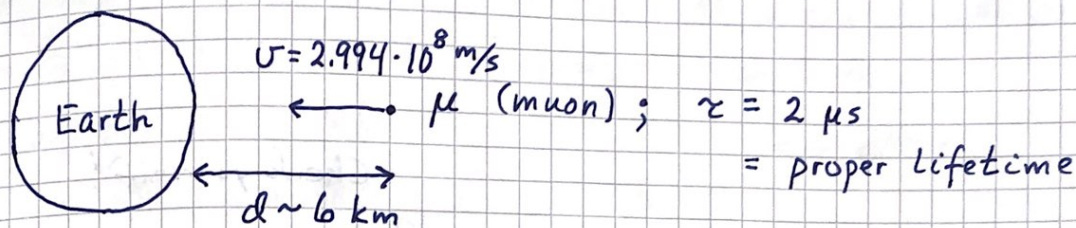
is longer than the time measured on a clock in S' ,

where the particle is at rest

Or: Moving clocks run slow.

Ex: Lifetime of muons in the atmosphere

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Nonrel.: $l = v \cdot \tau \approx 600 \text{ m} \ll d$

Rel.: $t = \gamma \tau \approx 32 \mu\text{s} = \text{lifetime measured on earth}$
 $l = v \cdot t \approx 9.6 \text{ km} > d$

⇒ Muons created in the atmosphere arrive on earth during their lifetime, in agreement with experiment.
 [Bailey et al, Physics Letters 55B, 420 (1975)]

Classification of 4-vectors:

space-like : $X_\mu X_\mu > 0$

time-like : $X_\mu X_\mu < 0$

light-like : $X_\mu X_\mu = 0$

Assume S' has speed v along $\hat{z}$ relative to S .

Consider events 1 and 2 described by $x_{1\mu}$ and $x_{2\mu}$ in S . The difference $\Delta x_\mu = x_{1\mu} - x_{2\mu}$ has norm

$$\Delta x_\mu \Delta x_\mu = (z_1 - z_2)^2 - c^2 (t_1 - t_2)^2$$

The time difference in S' is

$$t'_1 - t'_2 = \gamma \left[t_1 - t_2 - \frac{v}{c^2} (z_1 - z_2) \right] = \frac{\gamma}{c} \left[c(t_1 - t_2) - \frac{v}{c} (z_1 - z_2) \right]$$

space-like $\Delta x_\mu \Rightarrow \Delta x_\mu \Delta x_\mu > 0 \Rightarrow c < \left| \frac{z_1 - z_2}{t_1 - t_2} \right|$ (87)

$\Rightarrow$ 1 and 2 cannot be connected with a signal with speed c

$\Rightarrow$ the two events cannot influence each other

However, we can find a $v \leq c$ that gives

$$t'_1 - t'_2 = 0 \quad \text{or} \quad t'_1 - t'_2 \geq 0$$

I.e., before and after are not uniquely determined.

time-like $\Delta x_\mu \Rightarrow \Delta x_\mu \Delta x_\mu < 0 \Rightarrow c > \left| \frac{z_1 - z_2}{t_1 - t_2} \right|$

$\Rightarrow$ 1 and 2 can be connected with a signal with speed c

$\Rightarrow$ the two events can influence each other

Now, $t'_1 - t'_2$ and $t_1 - t_2$ have the same sign

I.e., before and after are uniquely determined

Causality: Cause before effect.

4-velocity: $u_\mu \equiv dx_\mu / d\tau = (dx_i, ic dt) / d\tau$

$$\Rightarrow u_i = dx_i / d\tau = \gamma dx_i / dt = \gamma v_i$$

$$u_4 = dx_4 / d\tau = ic dt / d\tau = ic \gamma$$

$$\Rightarrow u_\mu = \gamma (\vec{v}, ic)$$

Time-like:

$$u_\mu u_\mu = \gamma^2 v^2 - \gamma^2 c^2 = \frac{v^2 - c^2}{1 - v^2/c^2} = -c^2 < 0$$

4-current density:

The continuity equation is $\nabla \cdot \vec{j} + \partial \rho / \partial t = 0$

with $\vec{j}$ = current density, ρ = charge density

With $j_\mu = (\vec{j}, ic\rho)$: $\frac{\partial}{\partial x_\mu} j_\mu = 0$ or $\partial_\mu j_\mu = 0$

Due to length contraction, the charge density (88)
is larger for moving charges than for charges at rest:

$$\rho = \gamma \rho_0 > \rho_0$$

Hence,

$$j_\mu = (\vec{j}, ic\rho) = (\rho\vec{u}, ic\rho) = (\gamma\rho_0\vec{u}, ic\gamma\rho_0) = \rho_0 u_\mu$$

Let us show that the continuity eqn. is covariant.

We must show that $\partial/\partial x_\mu$ transforms like dx_μ under a LT. We have (from p.83) $dx_\mu = \underbrace{L_{\nu\mu}}_{=\partial x_\mu/\partial x'_\nu} dx'_\nu$

$$\Rightarrow \frac{\partial}{\partial x'_\nu} = \frac{\partial x_\mu}{\partial x'_\nu} \frac{\partial}{\partial x_\mu} = L_{\nu\mu} \frac{\partial}{\partial x_\mu}$$

$\Rightarrow$ Both $\partial/\partial x_\mu$ and $j_\mu = \rho_0 u_\mu$ transform as dx_μ under a LT

$\Rightarrow \partial_\mu j_\mu$ is a Lorentz invariant scalar

$\Rightarrow$ The cont. eqn. is covariant

Maxwell's equations and the 4-potential:

From 1.5; assuming vacuum ($\vec{P} = \vec{M} = 0$):

$$\vec{B} = \nabla \times \vec{A} \quad ; \quad \vec{E} = -\nabla\phi - \frac{\partial \vec{A}}{\partial t}$$

(using $\nabla \cdot \vec{B} = 0$ and $\nabla \times \vec{E} = -\partial \vec{B}/\partial t$)

$$\nabla \cdot \vec{E} = \rho/\epsilon_0 \Rightarrow \nabla^2 \phi + \frac{\partial}{\partial t} \nabla \cdot \vec{A} = -\rho/\epsilon_0$$

$$\nabla \times \vec{B} = \mu_0 \vec{j} + \epsilon_0 \mu_0 \partial \vec{E}/\partial t \quad ; \quad \epsilon_0 \mu_0 = 1/c^2$$

$$\Rightarrow \nabla^2 \vec{A} - \frac{1}{c^2} \frac{\partial^2 \vec{A}}{\partial t^2} - \nabla \left(\nabla \cdot \vec{A} + \frac{1}{c^2} \frac{\partial \phi}{\partial t} \right) = -\mu_0 \vec{j}$$

(using the identity $\nabla \times \nabla \times \vec{A} = \nabla(\nabla \cdot \vec{A}) - \nabla^2 \vec{A}$; see Assignment 2, Question 3)

Maxwell's 4 eqns. for $\vec{E}$ and $\vec{B}$ are reduced to 2 coupled eqns. for ϕ and $\vec{A}$; these may be decoupled since we have some freedom in our choice for ϕ and $\vec{A}$:

$$\vec{B} = \nabla \times \vec{A} \Rightarrow \vec{A} \rightarrow \vec{A} + \nabla \chi \text{ leaves } \vec{B} \text{ unchanged for arbitrary scalar } \chi, \text{ since } \nabla \times \nabla \chi = 0$$

Since $\vec{E} = -\nabla \phi - \partial \vec{A} / \partial t$, we must also let $\phi \rightarrow \phi - \partial \chi / \partial t$, to leave $\vec{E}$ unchanged.

This freedom to adjust ϕ and $\vec{A}$ is called gauge invariance.

Lorenz gauge (Ludvig Lorenz, Danish physicist, 1829-1891):

Choose a transformation that obeys

$$\nabla^2 \chi - \frac{1}{c^2} \partial^2 \chi / \partial t^2 = 0$$

Then, the Lorenz condition

$$\nabla \cdot \vec{A} + \frac{1}{c^2} \partial \phi / \partial t = 0$$

is gauge invariant,

$$\begin{aligned} & \nabla \cdot (\vec{A} + \nabla \chi) + \frac{1}{c^2} \frac{\partial}{\partial t} (\phi - \partial \chi / \partial t) \\ &= \underbrace{\nabla \cdot \vec{A} + \frac{1}{c^2} \partial \phi / \partial t}_{=0} + \underbrace{\nabla^2 \chi - \frac{1}{c^2} \partial^2 \chi / \partial t^2}_{=0} = 0 \end{aligned}$$

The Lorenz condition decouples the eqns. for $\vec{A}$ and ϕ :

$$\nabla^2 \vec{A} - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \vec{A} = -\mu_0 \vec{j}$$

$$\nabla^2 \phi - \frac{1}{c^2} \frac{\partial^2}{\partial t^2} \phi = -\rho / \epsilon_0$$

The 4-potential: $A_\mu = (\vec{A}, i\phi/c)$

$\Rightarrow$ Lorenz condition: $\partial_\mu A_\mu = 0$ [since $\frac{\partial}{\partial x_4} = \frac{\partial}{ic \partial t}$]

Eqn. for A_μ : $\square^2 A_\mu = -\mu_0 j_\mu$; $\square^2 \equiv \partial_\mu \partial_\mu = \nabla^2 - \frac{1}{c^2} \frac{\partial^2}{\partial t^2}$
= D'Alembert's operator

$\Rightarrow A_\mu$ transforms like $j_\mu = (\vec{j}, ic\rho)$ and u_μ and x_μ under a Lorentz transf. (Hendrik Lorentz, Dutch physicist, 1853-1928)

$\Rightarrow$ Maxwell's eqns. are covariant, i.e., consistent with SR!

7.4 Relativistic dynamics

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$N2$, $F = ma$, is invariant under a GT, not under a LT
 $\Rightarrow$ We must generalize Newtonian dynamics such that it

(a) becomes covariant, and

(b) reduces to $\frac{d}{dt}(mv_i) = F_i$ when $v_i \ll c$

Natural attempt:

$$\frac{d}{d\tau}(mu_\mu) = K_\mu \quad \text{with} \quad \begin{cases} m = \text{invariant mass} \\ \tau = \text{proper time} \\ u_\mu = 4\text{-velocity} \\ K_\mu = \text{Minkowski force} \end{cases}$$

Since $d\tau = dt/\gamma \rightarrow dt$ and $u_i = \gamma v_i \rightarrow v_i$ when $v \rightarrow 0$,
we must have $K_i \rightarrow F_i$ when $v \rightarrow 0$ (i.e. $v_i \ll c$)

Let us consider Maxwell's EM theory, since
it is covariant, with a known force $\vec{F} = q(\vec{E} + \vec{v} \times \vec{B})$.

The Lorentz force in terms of $\vec{A}$ and ϕ (i.e. A_μ):

$$\vec{F} = q \left(-\nabla\phi - \frac{\partial \vec{A}}{\partial t} + \vec{v} \times \nabla \times \vec{A} \right)$$

$$\stackrel{P.17}{\Rightarrow} F_i = -q \left[\partial_i (\phi - \vec{v} \cdot \vec{A}) + \frac{dA_i}{dt} \right]$$

$$\text{Since } u_\mu A_\mu = \gamma v_i A_i + \gamma ic \frac{i\phi}{c} = \gamma (\vec{v} \cdot \vec{A} - \phi)$$

and $d\tau = \gamma dt$, we have

$$\begin{aligned} F_i &= -q \left[-\frac{1}{\gamma} \partial_i (u_\mu A_\mu) + \frac{1}{\gamma} \frac{dA_i}{d\tau} \right] \\ &= \frac{q}{\gamma} \left[\partial_i (u_\mu A_\mu) - \frac{dA_i}{d\tau} \right] \end{aligned}$$

When $v \ll c$, $\gamma \rightarrow 1$, and we identify

$$K_i = q \left[\partial_i (u_\mu A_\mu) - \frac{dA_i}{d\tau} \right] = \gamma F_i$$

$$\text{and } K_\mu = q \left[\partial_\mu (u_\nu A_\nu) - \frac{dA_\mu}{d\tau} \right]$$

4-momentum: $p_\mu = m u_\mu = (m \gamma \vec{v}, i c \gamma m)$

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$$\Rightarrow K_\mu = \frac{dp_\mu}{d\tau}$$

$$\Rightarrow K_i = \gamma F_i = \frac{dp_i}{dt/\gamma} \Rightarrow \vec{F} = \frac{d\vec{p}}{dt} \quad \text{☺}$$

Energy and power:

$$u_\mu K_\mu = u_\mu \frac{d}{d\tau} (m u_\mu) = \frac{d}{d\tau} \left(\frac{1}{2} m u_\mu u_\mu \right) = 0$$

since $u_\mu u_\mu = -c^2$ (see p. 87)

$$\Rightarrow u_i K_i + u_4 K_4 = 0$$

$$\Rightarrow \gamma \vec{v} \cdot \gamma \vec{F} + i c \gamma \cdot K_4 = 0$$

$$\Rightarrow K_4 = \frac{i \gamma}{c} \vec{F} \cdot \vec{v} \Rightarrow K_\mu = \gamma \left(\vec{F}, \frac{i}{c} \vec{F} \cdot \vec{v} \right)$$

We combine $dp_4/d\tau = K_4$ with $p_4 = i c \gamma m$ and $K_4 = (i \gamma/c) \vec{F} \cdot \vec{v}$ and find

$$\vec{F} \cdot \vec{v} = \frac{d}{dt} (\gamma m c^2)$$

Since $\vec{F} \cdot \vec{v}$ is added power, we identify the total energy $E = \gamma m c^2$, and $p_4 = i \gamma m c = i E/c$

In summary: $\boxed{\vec{F} = d\vec{p}/dt, \vec{F} \cdot \vec{v} = dE/dt, p_\mu = (\vec{p}, i E/c)}$

- Rest energy ($v=0$): $E_0 = m c^2$
- Kinetic energy: $T = E - E_0 = (\gamma - 1) m c^2$
- Low speed: $\gamma \approx 1 + \frac{1}{2} \beta^2 \Rightarrow T \approx \frac{1}{2} m v^2$
- If the particle is in an external field, the potential energy V must be added to $\gamma m c^2$

Transformation of momentum and energy :

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Since $p_\mu = m dx_\mu / d\tau$, p_μ will transform as x_μ ,

$$p' = \mathbb{L} p$$

$$\Rightarrow p'_1 = p_1, p'_2 = p_2, p'_3 = \gamma p_3 + i\beta\gamma p_4 = \gamma(p_3 - \frac{v}{c^2} E)$$

$$p'_4 = -i\beta\gamma p_3 + \gamma p_4 = -i\beta\gamma p_3 + \gamma \frac{iE}{c} \quad \text{and} \quad p'_4 = \frac{iE'}{c}$$

$$\Rightarrow E' = \gamma(E - vp_3)$$

The invariant $p_\mu p_\mu$ in the particle's rest frame ($\vec{p}=0, \gamma=1$):

$$p_\mu p_\mu = -E_0^2/c^2 = -m^2 c^2$$

Hence, in general : $p^2 - E^2/c^2 = -m^2 c^2$

$$\Rightarrow \boxed{E^2 = p^2 c^2 + m^2 c^4}$$

Ex 1: Photons; $m=0 \Rightarrow E=pc$. Planck: $E=h\nu$

$$\Rightarrow pc = p\lambda\nu = h\nu \Rightarrow p = h/\lambda \quad (\text{de Broglie})$$

Ex 2: Fusion ; energy conservation and mass change

$$\begin{array}{c} \xrightarrow{v} \\ m(1) \end{array} \quad \begin{array}{c} \xleftarrow{-v} \\ m(2) \end{array} \Rightarrow \begin{array}{c} \bullet \\ M \end{array} \text{ at rest}$$

No external force $\Rightarrow$ conservation of $\vec{p} = \vec{p}_1 + \vec{p}_2 (=0)$

LT mixes E and $\vec{p} \Rightarrow$ energy also conserved

$\Rightarrow$ Total 4-momentum $P_\mu = (\vec{p}, iE/c)$ conserved

Conservation of P_4 : $\gamma mc^2 + \gamma mc^2 = Mc^2$

$$\Rightarrow M = 2\gamma m > 2m$$

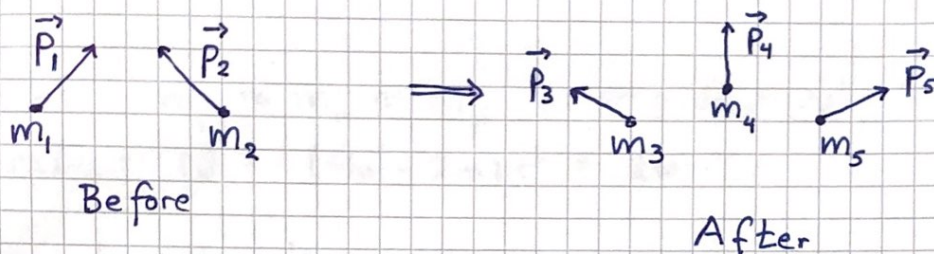
Kin. energy $\rightarrow$ Rest energy (mass): $\Delta M = M - 2m = 2m(\gamma - 1)$

7.5 Relativistic kinematics

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Relevance : Particle physics, High energy physics

Typical situation : 2 particles collide and form 2 or more new particles. Details of the interaction unknown, but conservation of momentum and energy, i.e., cons. of 4-momentum applies to "before" and "after".



$$P_\mu = P_{1\mu} + P_{2\mu}$$

$$P_\mu^r = P_{3\mu} + P_{4\mu} + P_{5\mu} \quad (+ \dots)$$

For each particle : $p_\mu = (\vec{p}, iE/c)$

Convenient inertial system :

The center-of-momentum system (COM) where

$$\vec{P} = \sum_j \vec{p}_j = 0$$

In the COM system we have

$$P_\mu P_\mu = P_\mu^r P_\mu^r = -M^2 c^2$$

which defines the equivalent mass M of the system.

Threshold energy = Minimum energy required for a reaction to take place ; corresponds to all reaction products at rest in the COM system after the collision.

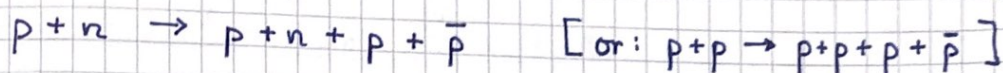
In that case, $M = \sum_r m_r$ = sum of mass of rxn. products

The Q-value of a reaction :

$$Q = \left[\sum_r m_r - (m_1 + m_2) \right] c^2 = \text{rest energy created in the reaction}$$

Ex: Antiproton production

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P.A.M. Dirac, Nobel Lecture 12.12.1933: Prediction of $\bar{p}$
O. Chamberlain, E. Segrè, C. Wiegand, T. Ypsilantis,
Phys. Rev. 100, 947 (1955): Observation of $\bar{p}$
[Chamberlain and Segrè, Nobel Prize 1959]

Masses: $m_p \approx m_n \approx m_{\bar{p}} \approx 938 \text{ MeV}/c^2 = m$

Q-value: $Q = (4m - 2m)c^2 = 2mc^2$

Calculate the threshold energy

(a) T_p with n at rest initially

(b) $T_p = T_n$ for head-on collision with $\vec{p}_p = -\vec{p}_n$

↳ Assignment 11

(a) We consider the covariant $P_\mu P_\mu$ in the lab system, to get T_p into the equation, and compare it with $P'_\mu P'_\mu$ after the collision in the COM system, where $M = 4m$. In other words,

$$P_\mu P_\mu = -M^2 c^2 = -16m^2 c^2$$

The left hand side is

$$P_\mu P_\mu = (P_{p\mu} + P_{n\mu}) \cdot (P_{p\mu} + P_{n\mu})$$

$$= P_{p\mu} P_{p\mu} + P_{n\mu} P_{n\mu} + 2 P_{p\mu} P_{n\mu}$$

$$= -m^2 c^2 - m^2 c^2 + 2 \left(\underbrace{\vec{p}_p \cdot \vec{p}_n}_0 + \frac{iE_p}{c} \cdot \frac{iE_n}{c} \right)$$

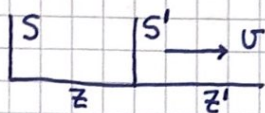
Here, $E_n = mc^2$ and $E_p = mc^2 + T_p$

$$\Rightarrow -4m^2 c^2 - 2mT_p = -16m^2 c^2 \Rightarrow T_p = 6mc^2 = \underline{\underline{5.57 \text{ GeV}}}$$

7.6 The electromagnetic field tensor and transformation of $\vec{E}$ and $\vec{B}$

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As usual, we assume:



Definition of EM field tensor:

$$F_{\mu\nu} = \frac{\partial A_\nu}{\partial x_\mu} - \frac{\partial A_\mu}{\partial x_\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

i.e. an antisymmetric tensor, $F_{\nu\mu} = -F_{\mu\nu}$

We have: $A_\mu = (\vec{A}, i\phi/c)$; $\vec{E} = -\frac{\partial \vec{A}}{\partial t} - \nabla\phi$; $\vec{B} = \nabla \times \vec{A}$

Hence:

$$F_{\mu\nu} = \begin{pmatrix} 0 & B_3 & -B_2 & -iE_1/c \\ -B_3 & 0 & B_1 & -iE_2/c \\ B_2 & -B_1 & 0 & -iE_3/c \\ iE_1/c & iE_2/c & iE_3/c & 0 \end{pmatrix}$$

Lorentz transformation of rank 2 tensor:

$$F'_{\mu\nu} = L_{\mu\alpha} L_{\nu\beta} F_{\alpha\beta}$$

with

$$L_{\mu\nu} = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & \gamma & i\beta\gamma \\ 0 & 0 & -i\beta\gamma & \gamma \end{pmatrix}$$

Now, it is straightforward to find $\vec{E}'$ and $\vec{B}'$ expressed in terms of $\vec{E}$, $\vec{B}$ and $\vec{v}$.

Let us derive a couple of components, then write down the remaining components, and finally consider the nonrelativistic limit $v \ll c$.

$$\mu=4, \nu=1:$$

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$$\frac{i}{c} E_1' = F_{41}' = L_{4\alpha} L_{1\beta} F_{\alpha\beta} = -i\beta\gamma B_2 + \gamma \cdot \frac{i}{c} E_1$$

$$\Rightarrow E_1' = \gamma (E_1 - v B_2)$$

$$\text{Assignment 11: } E_2' = \gamma (E_2 + v B_1); \quad E_3' = E_3$$

$$\text{If } v \ll c, \quad \gamma = 1, \text{ and } \vec{E}' = \vec{E} + \vec{v} \times \vec{B}$$

$$i=2, k=3:$$

$$\begin{aligned} B_1' = F_{23}' &= L_{2\alpha} L_{3\beta} F_{\alpha\beta} = L_{33} F_{23} + L_{34} F_{24} \\ &= \gamma B_1 + i\beta\gamma \cdot \left(-\frac{i}{c} E_2\right) = \gamma \left(B_1 + \frac{\beta}{c} E_2\right) \end{aligned}$$

$$\text{Assignment 11: } B_2' = \gamma \left(B_2 - \frac{\beta}{c} E_1\right); \quad B_3' = B_3$$

$$\text{If } v \ll c, \quad \gamma = 1, \text{ and } \vec{B}' = \vec{B} - \frac{1}{c^2} (\vec{v} \times \vec{E})$$